

Chapter 13: Universal Gravitation

Background/Summary:

In this unit, we will delve into different concepts surrounding gravity, orbits, gravitational potential energy, escape velocity, and more.

Major Topics, Vocabulary, and Formulae:

1. Newton's Law of Universal Gravitation

In the past, we have calculated the force of gravity only on Earth using the formula $F_g = mg$, but there is another formula for calculating F_g that takes into account the mass of the object (m_1), the mass of the body (m_2) exerting the force of gravity, the distance between the two objects (r), and the unit vector between the two (\hat{r}). We use this formula when the acceleration caused by gravity is not g .

Newton's Law of Universal Gravitation:

$$F_g = G \frac{m_1 m_2 \hat{r}}{r^2} \quad \text{where } G = 6.67 * 10^{-11} \frac{(N \cdot m^2)}{kg^2}$$

The "G" in this equation is the symbol for the universal gravitational constant, which, according to the equations sheet, equals. This law applies to all objects.

Note: when you are given the radius of an object from Earth, remember to include the radius of Earth ($r = 6.38 * 10^6 m$) in your total radius for the equation. We also sometimes call Earth's mass "M."

Fun fact: we can also use this equation to derive the acceleration of objects on earth (g) with the following derivation:

$$\begin{aligned} F_g &= -\frac{GMm}{r} \\ ma_g &= -\frac{GMm}{r} && \text{plug in } G \text{ and earth's} \\ a_g &= -\frac{GM}{r} = \frac{(6.672e-11)(5.98e24)}{(6.38e6)^2} && \text{mass and radius} \\ &= \boxed{-9.80 m/s^2} \end{aligned}$$

Another way you may see this equation is in calculating geosynchronous orbit, which refers to the orbit of an object matching Earth's rotation.

$$\begin{aligned} F_g &= F_{\text{centripetal}} && \text{use Newton's Universal Law of} \\ \frac{GMm}{r^2} &= \frac{mv^2}{r} && \text{Gravitation and } F_c = ma_c \\ &&& F_c = m \frac{v^2}{r} \\ \frac{GMm}{r^2} &= \frac{mv^2}{r} && \text{cancel masses and radius} \\ \frac{GM}{r} &= v^2 \\ \boxed{v} &= \sqrt{\frac{GM}{r}} && \text{for geosynchronous orbit} \end{aligned}$$

2. Gravitational Potential Energy

Another concept covered within this unit is gravitational potential energy. In previous

units, we used $\Delta U = U_f - U_i = -\Delta W = \int_{x_i}^{x_f} F_g \cdot dx$. However, as highlighted in the

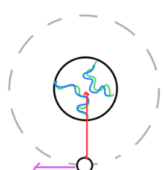
previous section, we need to use Newton's Law of Universal Gravitation when calculating F_g , and we would also need to use radius, not x value. The derivation for this equation is as follows:

$$\begin{aligned} \Delta U &= \int_{r_i}^{r_f} F_g \cdot dr && \text{plug in } F_g = -\frac{GMm}{r^2} \\ \Delta U &= \int_{r_i}^{r_f} -\frac{GMm}{r^2} \cdot dr && \text{pull the constants out of the integral} \\ \Delta U &= -GMm \int_{r_i}^{r_f} \frac{1}{r^2} dr && \text{solve the integral} \\ \Delta U &= -GMm \left(\frac{1}{r_f} - \frac{1}{r_i} \right) \\ &\rightarrow U_i = 0, \text{ and } r_i = \infty \\ \boxed{U} &= \boxed{-\frac{GMm}{r}} && \text{and when looking at a system with multiple masses,} \\ \boxed{U_{\text{total}}} &= \boxed{\sum U} \end{aligned}$$

3. Orbits and Energy

Recall that the energy of a system is $E_{\text{total}} = U + K = \frac{-GMm}{r} + \frac{1}{2}mv^2$. Within this unit, we can consider problems where there is a satellite, which is a body in orbit around another, that is "bound" in a circular orbit with a centripetal force provided by gravity. We can calculate the energy in the system using the following analysis:

Force analysis:



$$\begin{aligned} F_c &= F_g \\ \frac{r}{2} \cdot \frac{mv^2}{r} &= \frac{GMm}{r^2} \cdot \frac{r}{2} \\ \frac{1}{2}mv^2 &= \frac{GMm}{2r} \\ \underline{\underline{K}} &= \underline{\underline{-\frac{1}{2}U}} \end{aligned}$$

w want to find K, which is $\frac{1}{2}mv^2$, so we multiply both sides by $\frac{r}{2}$ to get K on the left

Back to the equation $E = U + K$, if we plug in $K = -\frac{1}{2}U$

$$\begin{aligned} E &= U + K \\ E &= U - \frac{1}{2}U = \boxed{\frac{1}{2}U} \text{ or } \boxed{-\frac{GMm}{2r}} \end{aligned}$$

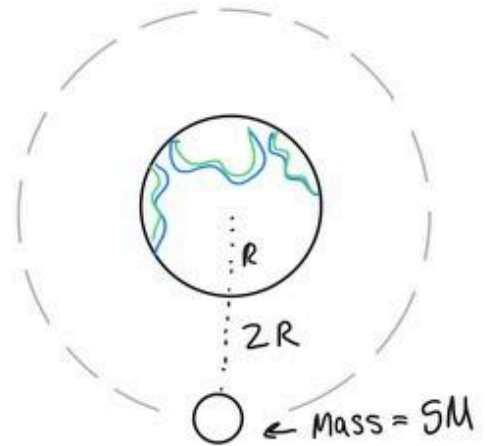
It makes sense that the energy of this system is negative because the gravitational potential energy of a system is negative when the radius $< \infty$.

Another context in which we could see energy analysis is when calculating the escape velocity of an object, meaning the velocity required for an object to completely leave earth's orbit.

$$\begin{aligned} K_i + U_i &= K_f + U_f \\ \frac{1}{2}mv_{\text{esc}}^2 + \frac{-GMm}{r} &= 0 + 0 \\ v_{\text{esc}} &= \sqrt{\frac{2GM}{r_i}} \end{aligned}$$

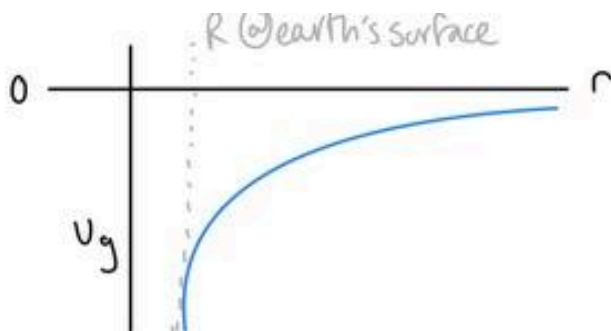
Example Problems/FRQs

1. Draw a graph of the gravitational potential energy between the Earth and the object as a function of the radius between the center of the Earth and the object, starting the graph with at the surface of the Earth. Provide a justification for this graph
2. A satellite orbits the earth with a time period $T=12$ hours. What is the orbital radius?
3. With the values displayed in the diagram to the right, in terms of G , M , m , and R , find...
 - a. Orbital speed
 - b. Gravitational potential energy at $r = 3R$
 - c. The escape velocity assuming that a similar object with mass $5M$ is launched from the surface of the earth



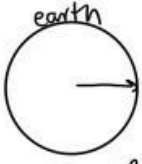
Example Problems/FRQ Answers (Part I)

Problem 1 Solutions:



From the equations sheet, $U_g = -\frac{Gdm}{r}$,
so the radius from earth's surface is
inversely related to the U_g . As r increases
to ∞ , the U approaches 0.

Problem 2 Solutions:



$m_{\text{earth}} = 5.84 \times 10^{24}$
 $r_{\text{earth}} = 6.37 \times 10^6 \text{ m}$
 $T = 12 \text{ hrs.} \cdot \frac{3600 \text{ s}}{1 \text{ hr}} = 4.32 \times 10^4 \text{ s}$
 $v = \frac{d}{t} = \frac{\text{circumference}}{\text{period}} = \frac{2\pi r}{T}$

this leaves us to force comparison so we can relate v and r

$$F_c = F_g \quad \rightarrow \text{cancel } m \text{ and } r$$

$$\frac{mv^2}{r} = \frac{GMm}{r^2}$$

$$v^2 = \frac{GM}{r} \quad \rightarrow \text{plug in } v = \frac{2\pi r}{T}$$

$$\left(\frac{2\pi r}{T}\right)^2 = \frac{GM}{r}$$

$$\frac{4\pi^2 r^2}{T^2} = \frac{GM}{r} \quad \rightarrow \text{isolate } r$$

$$r^3 = \frac{GMT^2}{4\pi^2}$$

$$r = \sqrt[3]{\frac{(6.672 \times 10^{-11})(5.84 \times 10^{24})(4.32 \times 10^4)^2}{4\pi^2}}$$

$$r = 2.64 \times 10^7 \text{ m}$$

→ yay! not orbiting inside the earth!

Problem 3 Solutions:



a) Finding orbital

$$F_g = F_c$$

$$\frac{GMm}{r^2} = \frac{mv^2}{r} \quad \rightarrow \text{plug in } R = R + 2R$$

$$\frac{GM}{2R} = v^2 \quad \rightarrow \quad v = \sqrt{\frac{GM}{3R}}$$

b) Gravitational potential energy

$$U = -\frac{GMm}{R} \quad \rightarrow \text{plug in values}$$

$$= -\frac{5}{3} \frac{GMm}{R}$$

c) escape velocity

$$K_i + U_i = K_f + U_f$$

$$\frac{1}{2}mv_i^2 + -\frac{GMm}{r} = 0 + 0$$

$$\frac{1}{2}mv_i^2 = \frac{GMm}{2R}$$

$$\frac{1}{2}v_i^2 = \frac{GM}{2R}$$

$$v_i = \sqrt{\frac{GM}{R}}$$

→ $v_f = 0$ when the rocket = out of earth's gravity as r increases, v_f approaches 0

→ get v on one side and cancel masses

→ isolate v